Space Weather with NASA's Van

with NASA's Van Allen Probes

Exploring the Extremes of Space Weather

Van Allen Probes

Mona Kessel, NASA HQ



Van Allen Space Weather Data

Space weather data is being generated and broadcast from the spacecraft 24/7 when not sending science data. The mission targets one part of the space weather chain: the very high energy electrons and ions magnetically trapped within Earth's radiation belts. The understanding gained by the Van Allen probes will enable us to better predict the response of the radiation belts to solar storms in the future, and thereby protect space assets in the near-Earth environment.

Outline

- Mission and Instrument Overview
- Capabilities for generating and broadcasting space weather data
 - Ground stations collecting the data
 - Data products
 - Collaboration with NOAA SWPC

Van Allen Mission Facts

Second Living With a Star Mission Launch August 30, 2012 Perigee: ~700 km altitude

Apogee ~5.5 Re geocentric altitude

Inclination ~10 degrees

Duration 2 years (+? expendables)

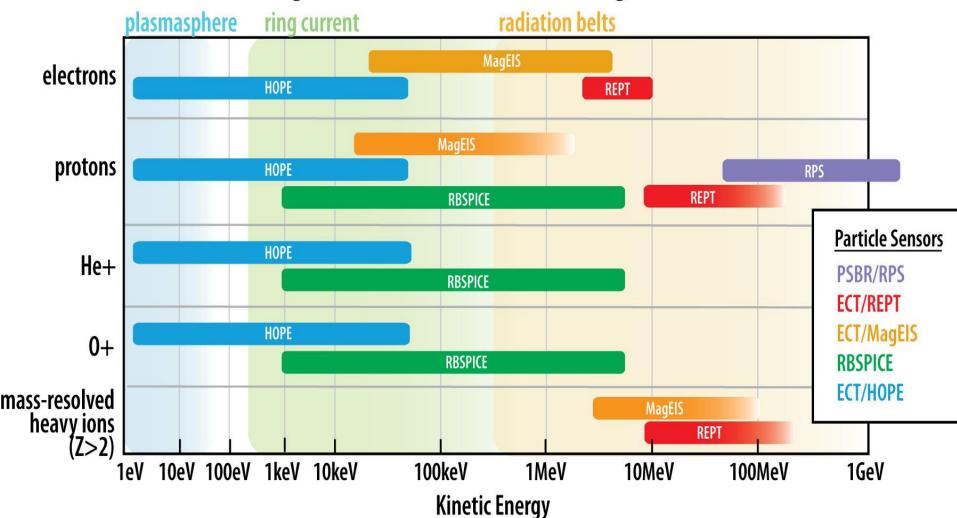
Mission Overview



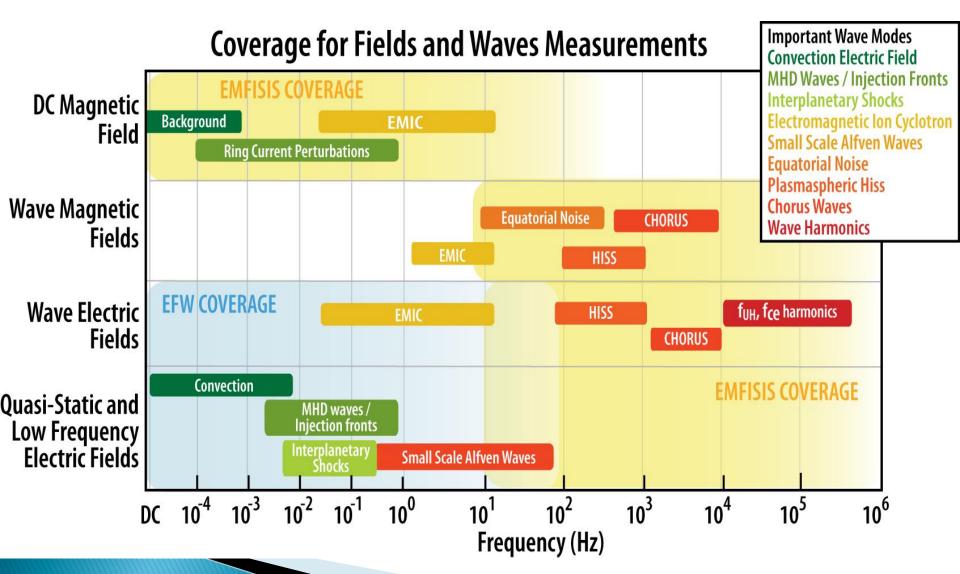
Provides understanding, *ideally to the point of predictability,* of how populations of relativistic electrons and penetrating ions in space form or change in response to variable inputs of energy from the Sun.

Particle Experiments

Coverage for Electron and Ion Pitch Angle Distributions

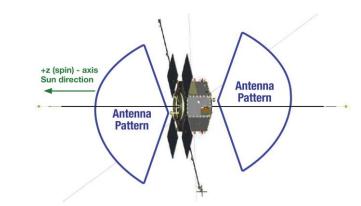


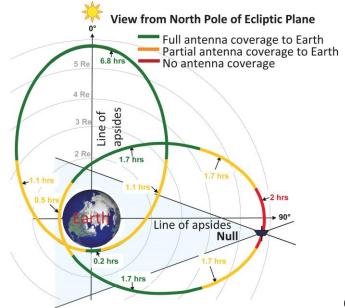
Field & Wave Experiments



Capabilities for generating and broadcasting space weather data

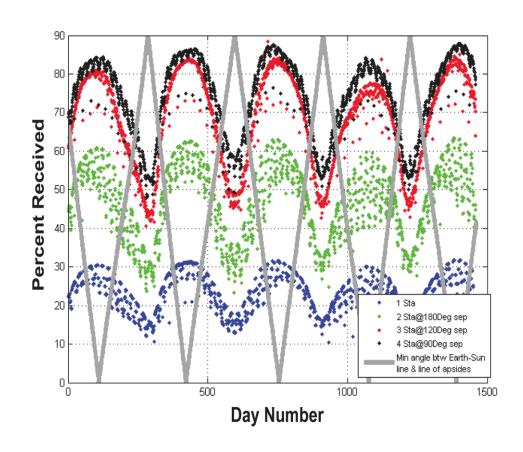
- Each satellite's sw broadcast ~21.5 hr/dy.
- Each instrument has a sw product.
- The 2 RF antennas are aligned with the spacecraft spin- and anti-spin-axes, coverage of ~140°.
- Both antennas are always active, only one has line-of-sight to Earth at any given time.
- Different regions of the orbit have communications downlinks that are robust, variable, or impossible.
- Stations near equator have best yeararound coverage.





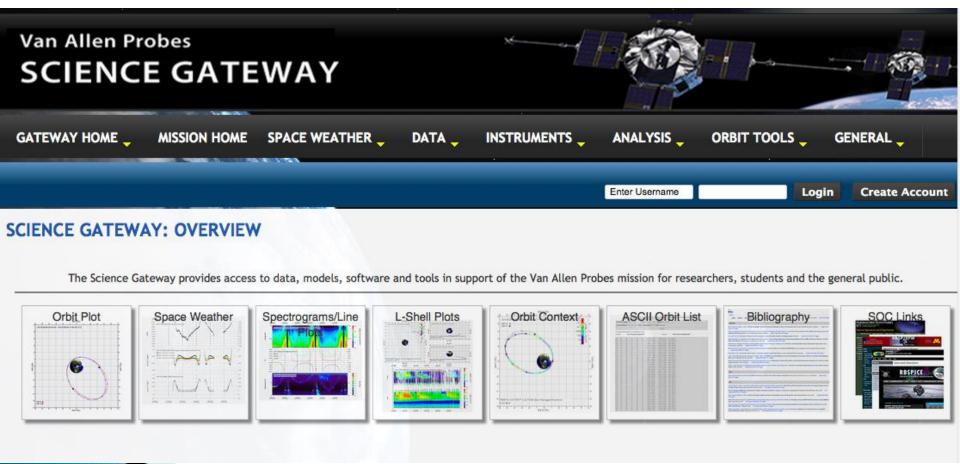
Space Weather Data Coverage

- Multiple stations at diverse longitudes can maximize potential spacecraft contact duration as a function of the number of ground stations.
- Agreements in place with 4 ground stations
 - Korea Astronomy and Space Science Institute (KASI)
 - Institute of Atmospheric Physics, Czech Republic
 - National Commission for Space Activities (CONAE), Argentina
 - National Institute of Space Research (INPE), Brazil



Van Allen Probes Gateway

The data are collected and processed at APL and then disseminated via Science Gateway tools

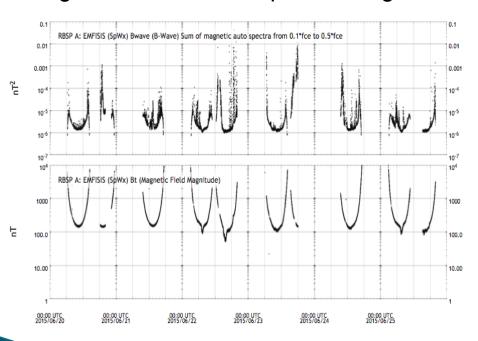


Space Weather Data Description

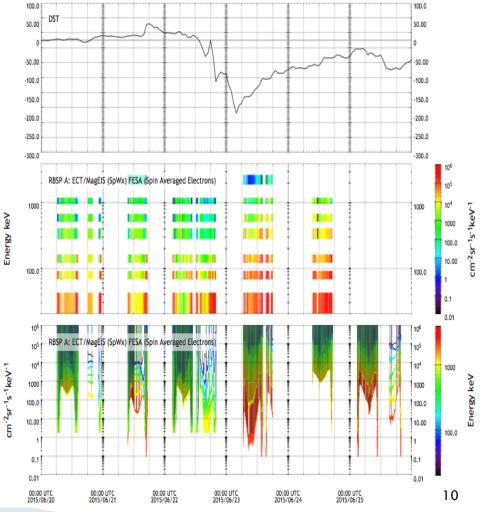
EMFISIS/MAG Vector Magnetic Field N/A EMFISIS/Waves VLF Wave Power N/A N/A Vector Electric Field N/A Spacecraft Potential N/A Electrons 24.54 eV, 281 eV, 10.9 keV, 42.9 keV** Protons 24.54 eV, 281 eV, 10.9 keV, 42.9 keV Oxygen Ions 24.54 eV, 281 eV, 10.9 keV, 42.9 keV Helium Ions 24.54 eV, 281 eV, 10.9 keV, 42.9 keV ECT/MagEIS Energetic Electrons 30 keV, 60 keV, 100 keV, 300 keV, 600 keV, 1 MeV ECT/REPT Very Energetic Electrons 2 MeV, 5 MeV, 10 MeV	strument	Measurement	Energy
Vector Electric Field N/A	WEISIS/MAG	Vector Magnetic Field	N/A
Spacecraft Potential N/A	MFISIS/Waves	VLF Wave Power	N/A
Spacecraft Potential N/A	EFW	Vector Electric Field	N/A
Protons 24.54 eV, 281 eV, 10.9 keV, 42.9 keV		Spacecraft Potential	N/A
Daygen lone 24.54 eV, 281 eV, 10.9 keV, 42.9 keV	ECT/HOPE	Electrons	24.54 eV, 281 eV, 10.9 keV, 42.9 keV**
Oxygen Ions 24.54 eV, 281 eV, 10,9 keV, 42.9 keV Helium Ions 24.54 eV, 281 eV, 10.9 keV, 42.9 keV Energetic Electrons 30 keV, 60 keV, 100 keV, 300 keV, 600 keV, 1 MeV Energetic Protons 1 MeV Very Energetic Electrons 2 MeV, 5 MeV, 10 MeV		Protons	24.54 eV, 281 eV, 10.9 keV, 42.9 keV
ECT/MagEIS Energetic Electrons 20 keV, 60 keV, 100 keV, 300 keV, 600 keV, 1 MeV Energetic Protons 1 MeV Very Energetic Electrons 2 MeV, 5 MeV, 10 MeV		Oxygen Ions	24.54 eV, 281 eV, 10,9 keV, 42.9 keV
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Very Energetic Electrons 2 MeV, 5 MeV, 10 MeV		Energetic Protons	1 MeV
ECT/KEPT	ECT/REPT	Very Energetic Electrons	2 MeV, 5 MeV, 10 MeV
Energetic Protons >20 MeV, >50 MeV, >70 MeV		Energetic Protons	>20 MeV, >50 MeV, >70 MeV
RBSPICE Energetic Protons 50 keV, 100 keV, 150 keV, 300 keV, 1 MeV, 10 MeV	BSPICE	Energetic Protons	50 keV, 100 keV, 150 keV, 300 keV, 1 MeV, 10 MeV
Energetic Protons >50 MeV, >400 MeV	PSBR/RPS	Energetic Protons	>50 MeV, >400 MeV
Dosimeter Data Linear and Log outputs (Volts)		Dosimeter Data	Linear and Log outputs (Volts)

Space Weather Example Data Plots

Magnetic Field – wave power, magnitude

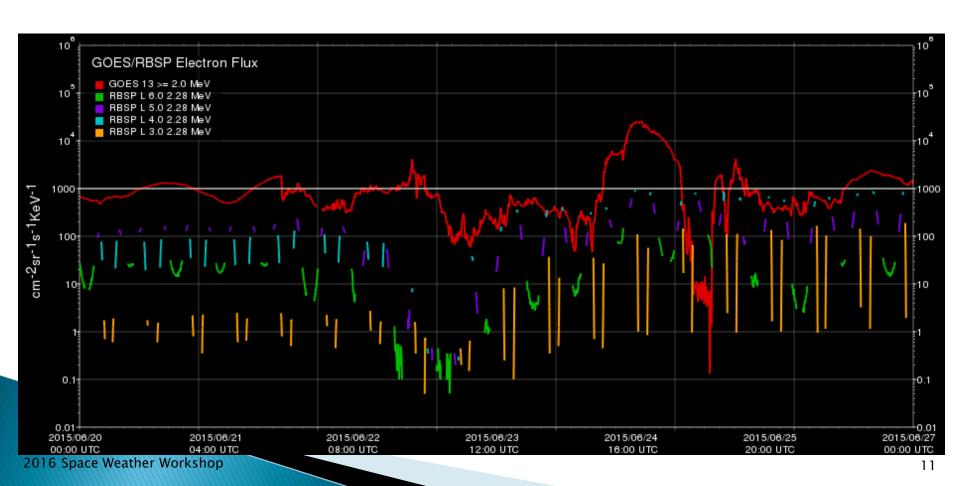


Dst, MagEIS spin avg electrons (2 views)



Van Allen Probes near-real-time data products soon to be re-broadcast by NOAA SWPC

- GOES characterizes radiation environment at GEO
- Van Allen Probes can fill in the L shells inside GEO

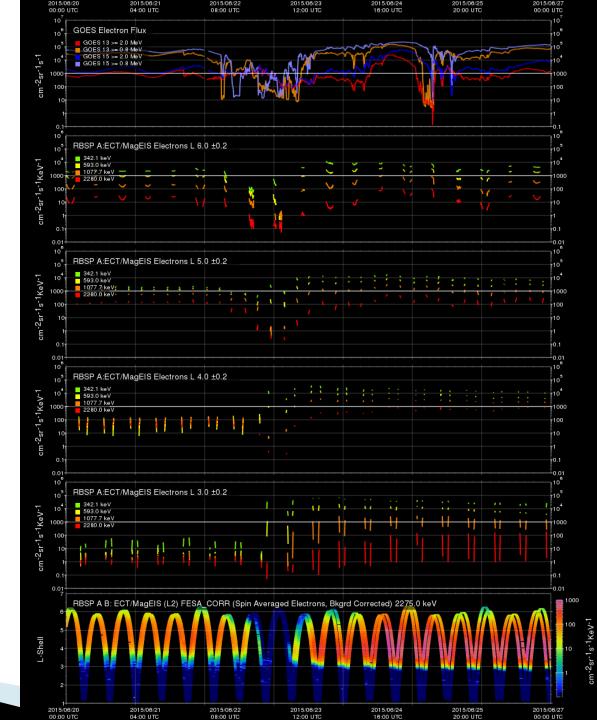


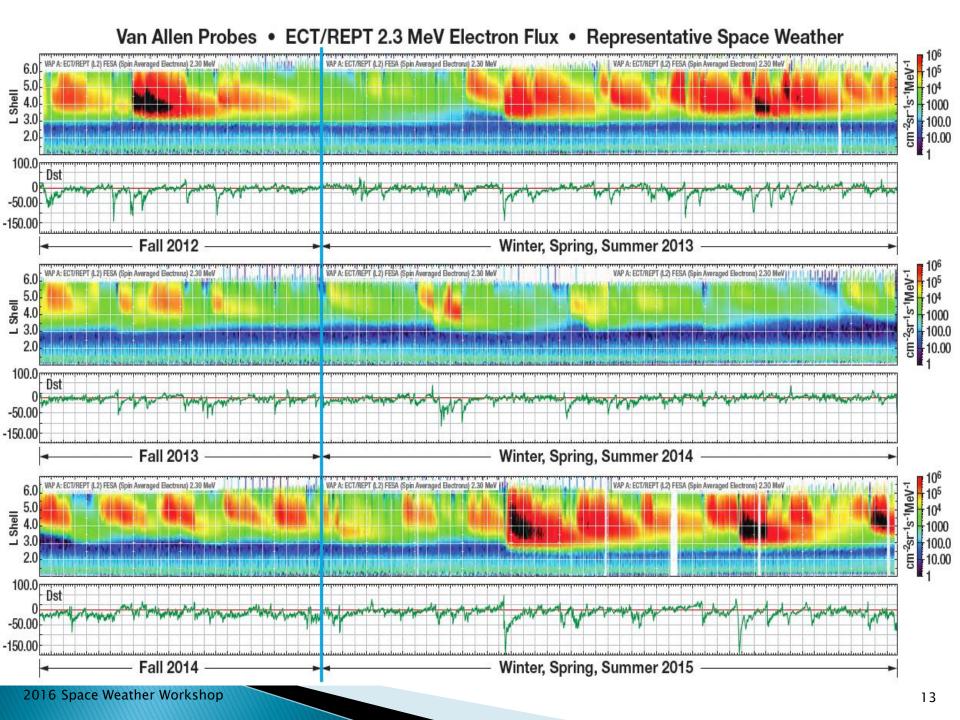
Lower energy electrons also show space weather effects

- 340 keV
- 590 keV
- 1.1 MeV

L shells from 6 to 3.

Van Allen Probes mission and NOAA SWPC are in process of finalizing the best inputs for forecasters.





Concluding Remarks

- The Van Allen probes continuously broadcast space weather data, except during prime science download and maneuvers. These data were selected to monitor the state of the radiation belts and will be used by NOAA SWPC to better space weather nowcasts.
- Currently four international partners have agreed to download this data and make it available for space weather data products. NASA is actively pursuing other ground station partners to fill in data gaps.
- The Van Allen probes have been designed to operate throughout the worst conditions expected in the hazardous radiation belt environment. The mission is significantly enhancing our understanding of radiation belt dynamics with changing solar wind conditions.
- The Van Allen probes will enable the prediction of extreme and dynamic space conditions, and will provide the understanding needed to design satellites to survive in space for future missions.